

$\mathcal{N} = 4$ supergravity amplitudes

Pierre Vanhove



Workshop on Black Holes in Supergravity and M/Superstring Theory

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based on

[arXiv:1202.3692] and [arXiv:1208.1255]

with Piotr Tourkine

Motivations

There are intriguing relations between supergravity and super-Yang-Mills amplitudes:

- ▶ Color-kinematic duality [see talk by Bern at Nicolai-fest]
- ▶ structure in the $1/N$ -expansion

These give some support to the relation that SUGRA field theory amplitudes can be presented as the square of SYM amplitudes

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At tree-level the relation $SUGRA \sim (SYM)^2$ can be understood simply by the use of the momentum kernel formalism [Kawai, Lewellen, Tye]

[Bjerrum-Bohr, Damgaard, Feng, Sondergaard], [Bjerrum-Bohr, Damgaard, Sondergaard, Vanhove]

Gravity \mathfrak{M}^{tree} and Yang-Mills \mathfrak{A}^{tree} amplitudes are build from the same data

$$\begin{aligned}\mathfrak{A}_n^{tree} &= A_n^{\text{spin } 1} \otimes S(k_i \cdot k_j) \otimes A_n^{\text{spin } 0} \\ \mathfrak{M}_n^{tree} &= A_n^{\text{spin } 1} \otimes S(k_i \cdot k_j) \otimes A_n^{\text{spin } 1}\end{aligned}$$

- $A_n^{\text{spin } s}$ are the color-stripped amplitudes for spin $s = 1, 0$
- there are in the kernel of the momentum kernel

$$\sum_{\sigma \in \mathfrak{S}_{n-2}} S[\beta | \sigma] A_n^{\text{spin } s}(1, \sigma(2, \dots, n-1), n) = 0, \quad \forall \beta \in \mathfrak{S}_{n-2}$$

- $S(k_i \cdot k_j)$ depends only the kinematics invariants

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This implies the [Bern-Carrasco-Johansson] parametrization of the amplitudes

$$\mathfrak{A}_n^{\text{tree}} = \sum_{\gamma \in \Gamma_n^{\text{tree}}} \frac{n_\gamma c^\gamma}{\prod_{\alpha \in \gamma} p_\alpha^2}; \quad \mathfrak{M}_n^{\text{tree}} = \sum_{\gamma \in \Gamma_n^{\text{tree}}} \frac{n_\gamma \tilde{n}^\gamma}{\prod_{\alpha \in \gamma} p_\alpha^2}$$

- ▶ Γ_n : ϕ^3 cubic graphs
- ▶ lorentz factor n_γ and color factors c^γ

$$SUGRA \sim (SYM)^2$$

At loop order this squaring relation require some care.

Since $1 \otimes 1 = 2 \oplus 0$, in the bosonic case propagate extra scalars

In supergravity theories the scalars parametrize a moduli space G/H and the invariant under the global symmetry G put strong constraints on the amplitudes

[[Green, Russo, Vanhove](#)], [[Elvang, Kiermaier, Freedman et al.](#)], [[Bossard, Stelle, Howe](#)]

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The $\mathcal{N} = 8$ supergravity multiplet is the tensor product of two $\mathcal{N} = 4$ SYM gauge multiplets

$$(2|_1, \frac{3}{2}|_8, 1|_{28}, \frac{1}{2}|_{56}, 0|_{\mathbf{70}})_{\mathcal{N}=8} = (1|_1, \frac{1}{2}|_4, 0|_{\mathbf{6}})_{\mathcal{N}=4} \otimes (1|_1, \frac{1}{2}|_{\bar{4}}, 0|_{\mathbf{6}})_{\mathcal{N}=4}$$

where we have indicated the $SU(8)$ and $SU(4)$ R-symmetry indices

The **70** scalars ϕ^i of the supergravity multiplet parametrize the coset $E_{7(7)}/(SU(8)/\mathbb{Z}_2)$.

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In the case $\mathcal{N} = 4$ SUGRA the gravity multiplet there is no unicity

$$\begin{aligned}(2|_{\mathbf{1}}, \frac{3}{2}|_{\mathbf{4}}, 1|_{\mathbf{6}}, \frac{1}{2}|_{\mathbf{4}}, 0|_{\mathbf{2}})_{\mathcal{N}=4} &= (1|_{\mathbf{1}}, \frac{1}{2}|_{\mathbf{4}}, 0|_{\mathbf{6}})_{\mathcal{N}=4} \otimes (1|_{\mathbf{1}})_{\mathcal{N}=0} \\ &= (1|_{\mathbf{1}}, \frac{1}{2}|_{\mathbf{2}}, 2 \times 0|_{\mathbf{1}})_{\mathcal{N}=2} \otimes (1|_{\mathbf{1}}, \frac{1}{2}|_{\bar{\mathbf{2}}}, 2 \times 0|_{\mathbf{1}})_{\mathcal{N}=2}\end{aligned}$$

The 2 scalars ϕ^i parametrize the coset $SU(1, 1)/U(1)$

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In this talk we will discuss the UV perturbative behaviour of $\mathcal{N} = 4$ and $\mathcal{N} = 8$ supergravity and discuss

- ▶ the role of supersymmetry in perturbative computation
- ▶ the role of non-perturbative duality symmetries in string theory

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There is no intellectual exercise that is not ultimately pointless
Jorge Luis Borges

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*There is no intellectual exercise that is not ultimately **stringy***
Jorge Luis Borges

Part I

$\mathcal{N} = 4$ SYM

Constraints from supersymmetry: $\mathcal{N} = 4$ SYM

- ▶ The case of $\mathcal{N} = 4$ super-Yang-Mills

$$\mathcal{S} = \frac{1}{g_{\text{YM}}^2} \int d^4x \text{tr}(F^2) + \dots$$

- ▶ Coupling constant dimensionless in $D = 4$

Half of supersymmetries are enough for finiteness of $\mathcal{N} = 4$ SYM in $D = 4$

[Mandelstam; Howe, Stelle, West; Brink, Lindgren, Nilsson]

- ▶ Four points amplitude behave as

$$\mathfrak{A}_{4;L}^{(D)} \sim \Lambda^{(D-4)L-4} t_8 F^4$$

But this is **not** enough for understanding the **correct** critical ultraviolet behaviour in the single trace sector, and double trace sector of the 4-point amplitudes in dimensions $4 < D \leqslant 10$

[Berkovits, Green, Russo, Vanhove], [Bern et al.]

Constraints from supersymmetry: $\mathcal{N} = 4$ SYM

$$[\mathfrak{A}_{4;L}^{(D)}] = \Lambda^{(D-4)L-4-2\gamma_L} \partial^{2\gamma_L} t_8 \text{tr}F^4 + \Lambda^{(D-4)L-4-2\beta_L} \partial^{2\beta_L} t_8 (\text{tr}F^2)^2$$

	L=1	L=2	L=3	L=4	L=5
$\partial^{2\gamma_L} t_8 \text{tr}(F^4)$	$D_c = 8$	$D_c = 7$	$D_c = 6$	$D_c = \frac{11}{2}$	$D_c = \frac{26}{5}$
	$\gamma_1 = 0$	$\gamma_2 = 1$	$\gamma_3 = 1$	$\gamma_4 = 1$	$\gamma_5 = 1$
$\partial^{2\beta_L} t_8 (\text{tr}F^2)^2$	$D_c = 8$	$D_c = 7$	$D_c = \frac{20}{3}$	$D_c = 6$	$D_c = \frac{28}{5}$
	$\beta_1 = 0$	$\beta_2 = 1$	$\beta_3 = 2$	$\beta_4 = 2$	$\beta_5 = 2$

- Some F-term are in $D < 10$

$$\partial^2 t_8 \text{tr}(F^4) \sim \int d^8 \theta \text{Tr}(W_\alpha^4)$$

$$\partial^4 t_8 (\text{tr}(F^2))^2 \sim \int d^{12} \theta (\text{Tr}(W_\alpha^2))^2$$

- Gaugino superfield $D_{(\beta} W_{\alpha)} = (\gamma^{mn})_{\alpha\beta} F_{mn}$

Constraints from supersymmetry: $\mathcal{N} = 4$ SYM

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- Some F-term are descendant of the Konishi operator $\text{tr}(\Phi \cdot \Phi)$ in $D < 10$

$$\begin{aligned} \partial^2 t_8 \text{tr}(F^4) &\sim \int d^{16}\theta \text{tr}(\Phi \cdot \Phi) \\ \partial^4 t_8 (\text{tr}(F^2))^2 &\sim \int d^{16}\theta (\text{tr}(\Phi \cdot \Phi))^2 \end{aligned}$$

- These operators are not protected from quantum corrections

Constraints from supersymmetry: $\mathcal{N} = 4$ SYM

$$[\mathfrak{A}_{4;L}^{(D)}] = \Lambda^{(D-4)L-4-2\gamma_L} \partial^{2\gamma_L} t_8 \text{tr}F^4 + \Lambda^{(D-4)L-4-2\beta_L} \partial^{2\beta_L} t_8 (\text{tr}F^2)^2$$

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	$\beta_1 = 0$	$\beta_2 = 1$	$\beta_3 = 2$	$\beta_4 = 2$	$\beta_5 = 2$

For $L \geq 4$ the UV divergence is dominated by the single trace term

single trace	$\Lambda^{(D-4)L-6} \partial^2 t_8 \text{tr}(F^4)$	$L \geq 2$
double trace	$\Lambda^{(D-4)L-8} \partial^4 t_8 (\text{tr}F^2)^2$	$L \geq 3$

- ▶ $\mathcal{N} = 3$ superspace explains the leading UV behaviour [[Howe, Stelle](#)]
- ▶ Confirmed by amplitude computation [[Bern, Dixon, Carrasco, Johansson, Roiban](#)] and in $D = 5$ by [[Bern, Douglas, ...; to appear](#)]

Part II

$\mathcal{N} = 8$ SUGRA

Behavior of supergravity amplitudes

$$\mathcal{S}^{gravity} = \frac{1}{2\kappa_{(D)}^2} \int d^D x \sqrt{-g} \mathcal{R}$$

- ▶ The gravitational coupling constant has dimension $[\kappa_{(D)}^2] = (length)^{D-2}$
- ▶ At L -loop gravity amplitudes have the superficial UV behaviour

$$[\mathfrak{M}_L^{(D)}] = \Lambda^{(D-2)L+2}$$

- ▶ The 4-graviton amplitudes factorize an $\mathcal{R}^4 = stu \mathfrak{M}_4^{\text{tree}}$ term

$$[\mathfrak{M}_L^{(D)}] = \Lambda^{(D-2)L-6-2\beta_L^N} \partial^2 \beta_L^N \mathcal{R}^4$$

- ▶ Critical dimension for UV divergences is

$$D \geq D_c = 2 + \frac{6 + 2\beta_L^N}{L}$$

The ultraviolet behaviour of $\mathcal{N} = 8$ supergravity

- For $L \leq 4$ the UV behaviour is

[Green, Russo, Vanhove; Bern et al.; Elvang et al.; Bossard et al.]

$$[\mathfrak{M}_{4;L}^{(D)}] \sim \Lambda^{(D-4)L-6} \partial^{2L} \mathcal{R}^4 \quad 2 \leq L \leq 4$$

The same as the one of $\mathcal{N} = 4$ SYM

- After 4-loop it is expected that only a $\partial^8 \mathcal{R}^4$ is factorized for $L \geq 4$

[Green, Russo, Vanhove; Vanhove; Green, Bjornsson]

$$[\mathfrak{M}_{4;L}^{(D)}] \sim \Lambda^{(D-2)L-14} \partial^8 \mathcal{R}^4 \quad L \geq 4$$

worse than the one of $\mathcal{N} = 4$ SYM

- At five-loop order the 4-point amplitude in
 - $\mathcal{N} = 4$ SYM divergences for $5 < 26/5 \leq D$
 - $\mathcal{N} = 8$ SUGRA divergences for $24/5 \leq D$

Double copy and UV behaviour

[Vanhove]

The double copy conjecture by BCJ state that the 4 point amplitude in $\mathcal{N} = 4$ SYM and $\mathcal{N} = 8$ SUGRA are related in a simple way

$$\begin{aligned}\mathfrak{A}_{4,L}^{(D)} &= g_{\text{YM}}^{2L+2} \sum_{\gamma \in \Gamma_4^L} \int \prod_{i=1}^L \frac{d^D \ell_i}{(2\pi)^D} \frac{1}{S_j} \frac{\textcolor{blue}{n}_\gamma c^\gamma}{\prod_{r=1}^{3L+1} p_r^2} \\ \mathfrak{M}_{4,L}^{(D)} &= \kappa_{(D)}^{2L+2} \sum_{\gamma \in \Gamma_4^L} \int \prod_{i=1}^L \frac{d^D \ell_i}{(2\pi)^D} \frac{1}{S_j} \frac{\textcolor{blue}{n}_\gamma \tilde{n}^\gamma}{\prod_{r=1}^{3L+1} p_r^2}.\end{aligned}$$

Double copy and UV behaviour

[Vanhove]

For large value of the loop momentum $\ell_i \sim \Lambda \rightarrow +\infty$ the amplitude behaves

$$\begin{aligned}\mathfrak{A}_{4,L}^{(D)} &\sim g_{\text{YM}}^{2L+2} \Lambda^{DL-2(3L+1)} n_\gamma^\infty \\ \mathfrak{M}_{4,L}^{(D)} &\sim \kappa_{(D)}^{2L+2} \Lambda^{DL-2(3L+1)} (n_\gamma^\infty)^2.\end{aligned}$$

The $\mathcal{N} = 4$ SYM leading UV behaviour is realized for numerator factor satisfying

- ▶ In the simple trace sector

$$n_\gamma^\infty|_{\text{single trace}} \sim \Lambda^{2L-4} \implies \mathfrak{A}_{4,L}^{(D)}|_{\text{single trace}} \sim \Lambda^{(D-4)L-6}$$

- ▶ In the double trace sector

$$n_\gamma^\infty|_{\text{double trace}} \sim \Lambda^{2L-6} \implies \mathfrak{A}_{4,L}^{(D)}|_{\text{double trace}} \sim \Lambda^{(D-4)L-8}$$

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This indicates that the leading UV behaviour of $\mathcal{N} = 8$ SUGRA is connected to the UV behaviour of SYM in the double trace sector

$$\mathfrak{M}_{4,L}^{(D)} \sim \Lambda^{DL-2(3L+1)} \underbrace{(n_\gamma^\infty|_{\text{double trace}})^2}_{=\Lambda^{2L-6}} = \Lambda^{(D-2)L-14}$$

Duality symmetry constraint on UV divergences

The UV behaviour can be derived by an application of the non-renormalisation theorems implied by the action of the extended symmetries and the duality group of $\mathcal{N} = 8$ SUGRA in various dimensions

$$\mathcal{S} = \frac{1}{\ell_D^{D-2}} \int d^D x \sqrt{-g} \left(\mathcal{R} + \ell_D^6 \mathcal{E}_{(0,0)}^{(D)} \mathcal{R}^4 + \ell_D^{10} \mathcal{E}_{(1,0)}^{(D)} D^4 \mathcal{R}^4 + \ell_D^{12} \mathcal{E}_{(0,1)}^{(D)} D^6 \mathcal{R}^4 + \dots \right)$$

The coupling $\mathcal{E}_{(0,0)}^{(D)}$ are function of the scalar fields ϕ^i in the coset $E_{n(n)}(\mathbb{R})/K_n$. In string theory there are automorphic form invariant under the action of $E_{n(n)}(\mathbb{Z})$

A UV divergence is a local counter-term $c_{UV} D^{2k} \mathcal{R}^4$ which is a part of the constant term of the string theory automorphic form $\mathcal{E}_{(p,q)}^{(D)}$ with $k = 2p + 3q$

Non-renormalisation theorems

For $D \geq 3$ [Green, Russo, Vanhove]

$$\left(\Delta^{(D)} - \frac{3(11-D)(D-8)}{D-2} \right) \mathcal{E}_{(0,0)}^{(D)} = 6\pi\delta_{D-8,0}$$

$$\left(\Delta^{(D)} - \frac{5(12-D)(D-7)}{D-2} \right) \mathcal{E}_{(1,0)}^{(D)} = 80\zeta(2)\delta_{D-7,0}$$

$$\left(\Delta^{(D)} - \frac{6(14-D)(D-6)}{D-2} \right) \mathcal{E}_{(0,1)}^{(D)} = -(\mathcal{E}_{(0,0)}^{(D)})^2 + \zeta(3)\delta_{D-6,0}$$

- The eigenvalues vanish in the critical dimensions for UV divergence $D^{2L}\mathcal{R}^4$ interactions

$$D_c = \begin{cases} 8 & \text{for } L=1 \\ 4+6/L & \text{for } 2 \leq L \leq 4 \end{cases}$$

- This has been confirmed using various field theory supersymmetry [Bossard, Stelle, Howe], and direct loop computation [Bern, Carrasco, Dixon, Johansson, Roiban], and continuous E_7 arguments [Elvang, Keirmaier, Freedman et al.]

Part III

$\mathcal{N} = 4$ SUGRA

Non-renormalisation theorems

We have seen that supersymmetry and dualities imply various non-renormalisation theorems for higher dimension operators.

- ▶ Type II compactifications on a torus ($\mathcal{N} = 8$ models) we had the $\beta_L^8 = L$ rule for $L \leq 4$ [Green, Russo, Vanhove; Berkovits]

$$[\mathfrak{M}_{4;L}^{(D)}] = \Lambda^{(D-4)L-6} D^{2L} \mathcal{R}^4, \quad \text{for } L \leq 4$$

- ▶ Heterotic compactification ($\mathcal{N} = 4$ models)
 - \mathcal{R}^4 is a $\frac{1}{2}$ -BPS coupling 1-loop exact in perturbation in string theory

[Bachas, Kiritsis; Bachas, Fabre, Kiritsis, Obers, Vanhove; Tourkine, Vanhove]

$$\beta_L^4 \geq 2 \quad \text{for } L \geq 2$$

Similarly to $\mathcal{N} = 8$ SUGRA these non-renormalisation theorems constraint the UV behaviour of $\mathcal{N} = 4$ supergravity

Constructing $\mathcal{N} = 4$ supergravity amplitudes

- ▶ The low-energy limit of superstring theory compactified on tori leads to pure $\mathcal{N} = 8$ supergravity.
- ▶ Obtaining pure $\mathcal{N} = 4$ supergravity from string theory is more tricky: generic string models have extra vector multiplets and there is no more a unique way to obtain the $\mathcal{N} = 4$ supergraviton multiplet

$$(2|_1, \frac{3}{2}|_4, 1|_6, \frac{1}{2}|_4, 0|_2)_{\mathcal{N}=4} = (1|_1, \frac{1}{2}|_4, 0|_6)_{\mathcal{N}=4} \otimes (1|_1)_{\mathcal{N}=0}$$

- ▶ or the product of two $\mathcal{N} = 2$ SYM gauge multiplets

$$(2|_1, \frac{3}{2}|_4, 1|_6, \frac{1}{2}|_4, 0|_2)_{\mathcal{N}=4} = (1|_1, \frac{1}{2}|_2, 2 \times 0|_1)_{\mathcal{N}=2} \otimes (1|_1, \frac{1}{2}|_{\bar{2}}, 2 \times 0|_1)_{\mathcal{N}=2}$$

Constructing $\mathcal{N} = 4$ supergravity amplitudes

- ▶ String theory constructions lead to models that has pure $\mathcal{N} = 4$ SUGRA coupled to $0 \leq n_v$ vector multiplet in four dimensions
- ▶ The string theory moduli space is (with $\Gamma \subset SL(2, \mathbb{Z})$)

$$\Gamma \backslash SU(1, 1, \mathbb{R}) / U(1) \times SO(6, n_v, \mathbb{Z}) \backslash SO(6, n_v, \mathbb{R}) / SO(6) \times SO(n_v)$$

To get pure $\mathcal{N} = 4$ supergravity we want to set $n_v = 0$ and decouple the string modes.

The string theory effective action is given by

$$\mathcal{S} = \frac{1}{\ell_4^2} \int d^4x e^{-\mathcal{R}} \left(\mathcal{R} + \ell_4^2 f(S, \bar{S}) R^2 + \ell_4^6 g(S, \bar{S}) \mathcal{R}^4 \right)$$

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To get pure $\mathcal{N} = 4$ supergravity we want to set $n_v = 0$ and decouple the string modes.

- ▶ For $(4, 0)$ models the complex scalar S of the supergravity multiplet is in the $SU(1, 1, \mathbb{R}) / U(1)$ factor
- ▶ For $(2, 2)$ models the scalar S parametrizes a $SU(1, 1, \mathbb{R}) / U(1) \subset SO(6, n_v, \mathbb{R}) / SO(6) \times SO(n_v)$
- ▶ these models are *non-perturbatively dual* in string theory

The predicted UV behaviour for $\mathcal{N} = 4$ supergravity

Taking the limit of one- and two-loop four-graviton amplitudes in string theory [Tourkine, Vanhove] match the supergravity expression for the supergraviton contribution obtained by [Bern et al., Dixon et al., Dunbar et al.]

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We find that the string theory non-renormalisation theorem applies to the pure $\mathcal{N} = 4$ supergravity limit and the critical UV behaviour of the 4-graviton amplitude is [Tourkine, Vanhove]

$\mathcal{N} = 4$ non-renormalisation theorems for \mathcal{R}^4 term $\beta_L^4 = 1$ for $L \geq 2$

$$[\mathfrak{M}_L^{(D)}] = \Lambda^{(D-2)L-8} \partial^2 \mathcal{R}^4 \quad \text{for } L \geq 2$$

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- The absence of UV divergence in $D = 4$ at $L = 3$ was obtained from direct field theory computation [Bern, Davies, Dennen, Huang]

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- The absence of UV divergence in $D = 5$ at $L = 2$ confirmed by direct field theory computation [Bern, Davies, Dennen, Huang (to appear)]

Double copy and UV behaviour

The double copy conjecture procedure for the $(4,0)$ construction instructs us to build the $\mathcal{N} = 4$ SUGRA amplitude from the product of numerators from $\mathcal{N} = 4$ SYM and pure YM

$$\mathfrak{M}_{4,L}^{(D)} = \kappa_{(D)}^{2L+2} \sum_{\gamma \in \Gamma_4^L} \int \prod_{i=1}^L \frac{d^D \ell_i}{(2\pi)^D} \frac{1}{S_j} \frac{n_\gamma|_{\mathcal{N}=4} \tilde{n}_\gamma^\gamma|_{YM}}{\prod_{r=1}^{3L+1} p_r^2}.$$

the $\mathcal{N} = 4$ supergravity UV behaviour

$$\mathfrak{M}_{4,L}^{(D)} \sim \Lambda^{DL-2(3L+1)} n_\gamma^\infty|_{\mathcal{N}=4 \text{ double trace}} \tilde{n}_\gamma^\infty|_{YM \text{ subleading}} = \Lambda^{(D-2)L-8}$$

is obtained from the subleading UV behaviour in pure YM and $\mathcal{N} = 4$ SYM

$$n_\gamma^\infty|_{YM \text{ leading}} \sim \Lambda^{2L+2}; \quad n_\gamma^\infty|_{YM \text{ subleading}} \sim \Lambda^{2L}$$

Duality symmetries of $\mathcal{N} = 4$ pure supergravity

- ▶ In [Bossard, Howe, Stelle, Vanhove] a candidate 3-loop counter-term was constructed using harmonic superspace

$$\int d^4x \int d^{12}\theta (\bar{\chi}\chi)^2 = \int d^4x e (\mathcal{R}^4 + \text{susy completion})$$

- ▶ This term is unique at this order if the duality symmetry $SU(1, 1)$ is exact since we showed that the duality invariant supervolume vanishes

$$\int d^4x \int d^{16}\theta E(x, \theta) = 0$$

- ▶ But there is no 3-loop divergences in $D = 4$ [Bern et al., Tourkine, Vanhove]
- ▶ It could be the off-shell properties of $\mathcal{N} = 4$ supergravity coupled to 6 vector multiplets make this coupling a protected F-term

Duality symmetries of $\mathcal{N} = 4$ pure supergravity

$\mathcal{N} = 4$ supergravity is special because of the $U(1)$ R-symmetry anomaly

[Marcus].

Therefore the $SU(1, 1)$ duality symmetry is broken in perturbation and more general functions of the axion-dilaton $S \in SU(1, 1)/U(1)$ are allowed for the coefficient of the counter-term to the effective action

$$\int d^4x \int d^{16}\theta E(x, \theta) h(\$) = \int d^4x e g(S, \bar{S}) \mathcal{R}^4 + \text{susy completion}$$

In string theory $g(S, \bar{S})$ is a modular form invariance under $\Gamma \subset SL(2, \mathbb{Z})$

- ▶ The absence of three loop divergences shows that $\Delta_S g(S, \bar{S}) \neq 0$ where $\Delta_S = 4(\text{Im } S)^2 \partial_S \bar{\partial}_{\bar{S}}$

Outlook

- ▶ Using string theory we have put constraints on the possible counter-terms for UV divergences of four gravitons amplitudes in $\mathcal{N} = 4$ and $\mathcal{N} = 8$ supergravity
- ▶ Absence of UV divergences is explained by fermionic zero mode saturation, which is a manifestation of supersymmetry
- ▶ In the gravity sector $\mathcal{N} = 4$ amplitudes are the same in the $(4, 0)$ and $(2, 2)$ construction: could be good to be tested from the double copy construction
- ▶ At the quantum level $\mathcal{N} = 4$ supergravity is ambiguous because of the anomaly. The ambiguity is seen in the coupling with the scalar fields

